



# Receptor modeling of near-roadway aerosol mass spectrometer data in Las Vegas, Nevada, with EPA PMF

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**Abstract.** Ambient non-refractory PM<sub>1</sub> aerosol particles were measured with an Aerodyne High Resolution Time-of-Flight Aerosol Mass Spectrometer (HR-AMS) at an elementary school 18 m from the US 95 freeway soundwall in Las Vegas, Nevada, during January 2008. Additional collocated continuous measurements of black carbon (BC), carbon monoxide (CO), nitrogen oxides (NO<sub>x</sub>), and meteorological data were collected. The US Environmental Protection Agency's (EPA) positive matrix factorization (PMF) data analysis tool was used to apportion organic matter (OM) as measured by HR-AMS, and rotational tools in EPA PMF were used to better characterize the solution space and pull resolved factors toward known source profiles. Three- to six-factor solutions were resolved. The four-factor solution was the most interpretable, with the typical AMS PMF factors of hydrocarbon-like organic aerosol (HOA), low-volatility oxygenated organic aerosol (LV-OOA), biomass burning organic aerosol (BBOA), and semi-volatile oxygenated organic aerosol (SV-OOA). When the measurement site was downwind of the freeway, HOA composed about half the OM, with SV-OOA and LV-OOA accounting for the rest. Attempts to pull the PMF factor profiles toward source profiles were successful but did not qualitatively change the results, indicating that these factors are very stable. Oblique edges were present in G-space plots, suggesting that the obtained rotation may not be the most plausible one. Since solutions found by pulling the profiles or using  $F_{\text{peak}}$  retained these oblique edges, there appears to be little rotational freedom in the base solution. On average, HOA made up 26 % of the OM, while LV-OOA was highest in the afternoon and accounted for 26 % of the OM. BBOA occurred in the evening hours, was predominantly from the residential area to the north, and

on average constituted 12 % of the OM; SV-OOA accounted for the remaining third of the OM. Use of the pulling techniques available in EPA PMF and ME-2 suggested that the four-factor solution was very stable.

## 1 Introduction

A number of studies have demonstrated the significant health impacts of air pollution such as particulate matter of aerodynamic diameter less than 2.5 microns (PM<sub>2.5</sub>), including increased asthma rates, detrimental fetal development during pregnancy, and decreased lung capacity (Brunekreef et al., 1997; McDonald et al., 2004; Dockery et al., 1993; Dockery and Stone, 2007). In particular, recent literature has demonstrated that adverse health effects are evident in subjects exposed to ambient air near major roadways (Edwards et al., 1994; Nitta et al., 1993; Kunzli et al., 2000; Hoek et al., 2002; Finkelstein et al., 2004; Kim et al., 2004b). A wide body of literature of previous near-roadway air pollution studies (Zhu et al., 2002; Zhang et al., 2004a; Phuleria et al., 2007; Ntziachristos et al., 2007; Fruin et al., 2008) has demonstrated that, for a number of pollutants such as black carbon (BC), carbon monoxide (CO), and ultrafine particles, concentrations are very high next to the freeway but rapidly decrease to background urban levels after a few hundred meters under persistent winds. In pre-sunrise hours and low wind conditions, the influence of freeway emissions can reach even further (Hu et al., 2009).

Organic matter (OM) is a large and important part of PM<sub>2.5</sub>, particularly in near-roadway environments, and is often the largest component of PM in urban areas in the

western United States (Phuleria et al., 2007; Riddle et al., 2008; Minguillon et al., 2008). OM is a complicated mixture of thousands of individual molecules and is a combination of both primary particulate emissions and secondary aerosol formed from gaseous precursors. It is a major component of vehicular exhaust emissions and, in addition to being a large part of PM<sub>2.5</sub> mass, also includes polycyclic aromatic hydrocarbons (PAHs), which are carcinogenic (Larsen and Baker, 2003; Lobscheid and McKone, 2004; Adonis et al., 2003; Flowers et al., 2002).

The composition of OM has been analyzed under a number of different methods, including molecular marker analysis of aerosol filter samples and real-time analysis with instruments such as the Aerodyne Aerosol Mass Spectrometer (AMS) and its more recent version, the High Resolution Time-of-Flight Aerosol Mass Spectrometer (HR-AMS) (Drewnick et al., 2005; Jimenez et al., 2003; Zhang et al., 2005a; Allan et al., 2004, 2003a, 2003b; DeCarlo et al., 2006). With the AMS, individual molecular marker compounds are not typically quantified, but with this loss of molecular specificity we gain high time resolution and a more complete representation of the full organic fraction. Rather than individual molecules, specific groups of mass-to-charge ratio fragments ( $m/z$ ) can be used to identify differences between less oxidized, hydrocarbon-like organic aerosol (HOA) and oxygenated organic aerosol (OOA). Some examples include data from Pittsburgh, Pennsylvania (Zhang et al., 2004b, 2005a, b), Riverside, California (Docherty et al., 2008), Zurich (Lanz et al., 2008a, 2007), Mexico City (Dzepina et al., 2007; Johnson et al., 2008; Aiken et al., 2008), and Fraser Valley, British Columbia (Alfarra et al., 2007, 2004).

Laboratory and field studies have found that  $m/z$  44 (i.e., CO<sub>2</sub><sup>+</sup>) is a major fragment when long range transport is important and during periods of active photochemistry and that  $m/z$  44 is representative of OOA. In prior field studies in Pittsburgh, British Columbia, and elsewhere (Zhang et al., 2004b, 2005a, b; Alfarra et al., 2007, 2004), saturated hydrocarbon fragments such as  $m/z$  57 (i.e., C<sub>4</sub>H<sub>9</sub><sup>+</sup>) were found to be typical of HOA.

AMS data have, in recent years, been further evaluated with positive matrix factorization (PMF) or other mathematical methods (Zhang et al., 2004b, 2005a, b) to decompose the mass spectra and quantify the amount of OOA and HOA impacting a monitoring site. Source studies of primary emissions—such as diesel exhaust in a vehicle-chasing experiment (Canagaratna et al., 2004), source profile acquisition (Mohr et al., 2009), and reaction chamber secondary OA (SOA) (Alfarra et al., 2006)—show similarities to spectra from ambient aerosol. These two types of primary OA (POA) and SOA profiles resemble those observed in the ambient air under conditions more conducive to primary (high  $m/z$  57 concentrations) and secondary (high  $m/z$  44 concentrations) influences, respectively. In addition, two types of OOA spectra have been observed in Riverside and elsewhere

(Docherty et al., 2008; Huffman et al., 2009; Jimenez et al., 2009), one with more “low volatility” fragments (LV-OOA) including  $m/z$  44, but also one with significant mass from other, less oxidized fragments. This spectrum has also been observed in chamber experiments with diesel exhaust and resembles semi-volatile OA with some oxidized OA, termed SV-OOA.

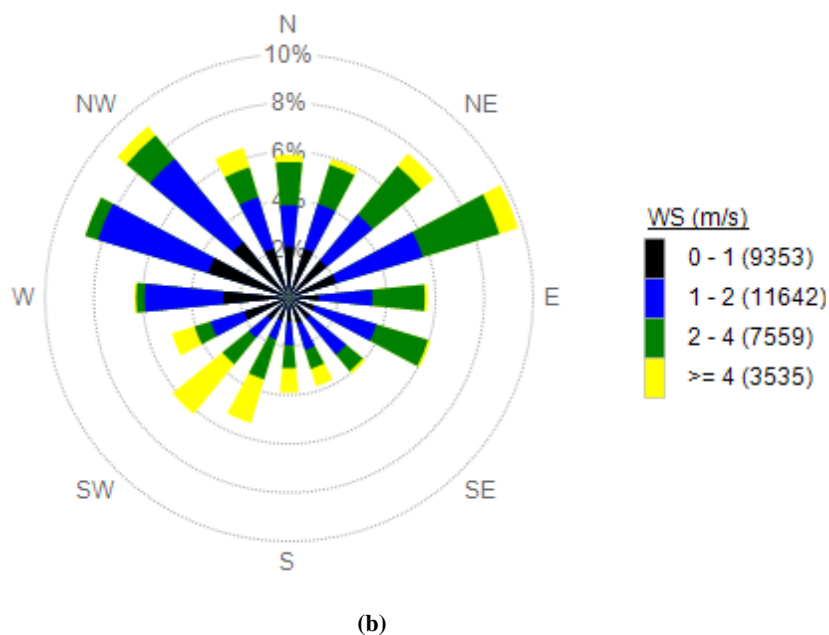
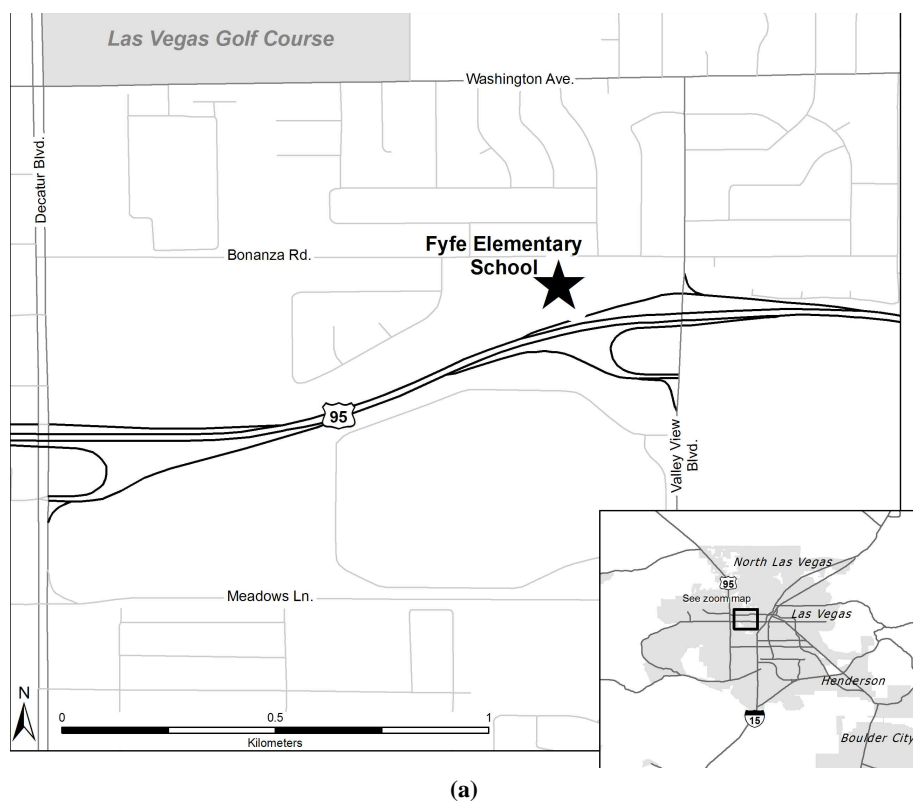
## 2 Methods

### 2.1 Aerosol and gaseous measurements

Measurements were made outdoors next to a classroom and play yard at Fyfe Elementary School, directly adjacent to and 18 m from the US 95 soundwall (Fig. 1). In 2007, the annual average daily traffic (AADT) was between 189 000 and 201 000 vehicles on the stretch of US 95 near Fyfe Elementary School. The wind sector from 90 to 250 degrees encompasses the adjacent freeway. CO, NO<sub>x</sub>, Aethalometer BC, wind speed, and wind direction were measured continuously, yielding 5-min averages. A two-channel (370 nm and 880 nm) Magee Scientific Aethalometer was used to measure BC in 5-min intervals. Aerosol was collected on a glass fiber tape, passing through a Harvard impactor with a size-cut of 2.5 microns. Raw data were post-processed with the Washington University Air Quality Lab AethDataMasher Version 6.0e to format date-time stamps and perform data validation. A Thermo Scientific 42i NO/NO<sub>2</sub>/NO<sub>x</sub> analyzer and a Thermo Scientific 48i CO analyzer provided 5-min NO<sub>x</sub> and CO data. Zero-checks and span-checks were performed nightly; data were zero-corrected if the zero-check was greater than 5 % of expected. Continuous gaseous instruments were calibrated at setup, take down, and quarterly, plus on an as-needed basis via remote control over the internet. Five-minute data were visually reviewed daily and after the study for additional QC. Wind speed and direction were measured with an RM Young AQ 5305-L at 1-min intervals, and were used to calculate vector-averaged 5-min average data.

### 2.2 HR-AMS description and data processing

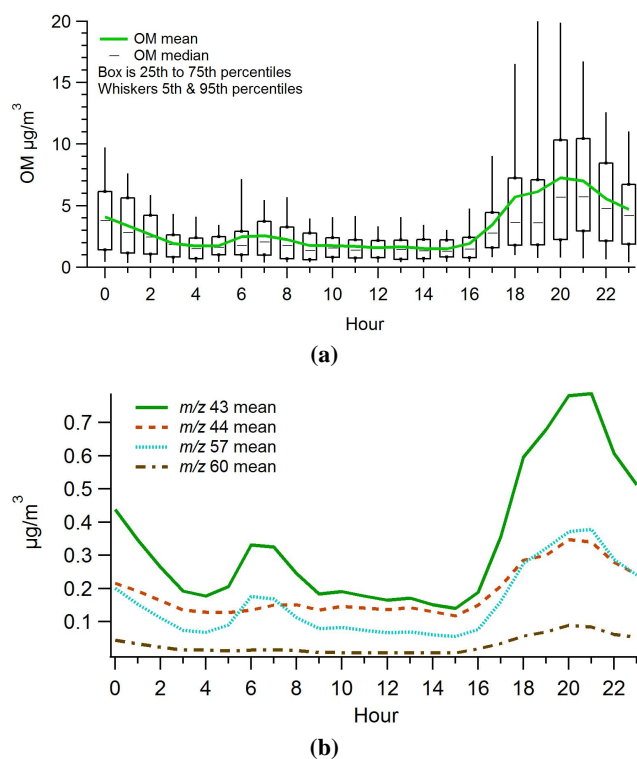
The operation of the Aerodyne HR-AMS has been described in detail elsewhere (DeCarlo et al., 2006; Drewnick et al., 2005; Jayne et al., 2000; Jimenez et al., 2003, 2009). Ambient air is drawn through a URG cyclone ( $D_{50} = 2.5\mu\text{m}$ , 3 lpm) and is sampled with 2-min time resolution through a critical orifice into an aerodynamic lens, creating a narrow particle beam, with a 50 % efficiency of 1 $\mu\text{m}$  particles, so that essentially PM<sub>1</sub> is measured (Sun et al., 2009; Canagaratna et al., 2007). The particles are accelerated in the supersonic expansion of gas molecules into a vacuum at the end of the lens. Particles are collected by inertial impaction and non-refractory species such as nitrate, sulfate, ammonium, chloride, potassium and OM are thermally vaporized.



**Fig. 1.** (a) location of Fyfe Elementary School in Las Vegas, Nevada; and (b) wind rose for January 2008 (5-min averages).

Vaporized gases undergo electron impact ionization and the charged fragments enter the ToF-MS region, where they are separated by mass-to-charge ratio ( $m/z$ ). After correction for ambient gases such as  $N_2$ , mass spectra are analyzed for each 2-min averaged sample, and the sum of organic aerosol

peaks is used to calculate total OM. AMS data were processed and analyzed using the standard AMS analysis software, Squirrel version 1.48, implemented with Wavemetric's Igor Pro (version 6.12). Concentration and uncertainty data for PMF were regenerated using Squirrel v 1.51. Uncertainty



**Fig. 2.** (a) Box plot of AMS OM ( $\mu\text{g m}^{-3}$ ) by hour, and (b) average concentrations ( $\mu\text{g m}^{-3}$ ) of selected  $m/z$  (43, 44, 57, and 60) by hour.

estimates are based on counting statistics methods and generated from Squirrel (Allan et al., 2003b). For a given AMS signal, the error estimate is:

$$S(j) \equiv \alpha \sqrt{\frac{[I(j_o) + I(j_b)]}{t(s)}} \quad (1)$$

where the ion signal is  $j$ , the signal when the beam is open is ( $I(j_o)$ ), and the signal when it is blocked is ( $I(j_b)$ ) over the sampling time  $t(s)$ , with a distribution factor  $\alpha$ .

### 2.3 Positive Matrix Factorization

PMF is a multivariate factor analysis tool (Paatero and Hopke 2009; Paatero, 1997; Paatero and Tapper, 1994, 1993) that has been applied to a wide range of data, including 24-hr speciated  $\text{PM}_{2.5}$  data, size-resolved aerosol data, deposition data, air toxics data, volatile organic compound (VOC) data (Kim et al., 2003; Kim and Hopke, 2004; Kim et al., 2004a; Polissar et al., 2001; Poirot et al., 2001; Brown and Hafner, 2003; Hopke, 2003), and more recently to AMS data sets (Lanz et al., 2008a; Docherty et al., 2008; Ulbrich et al., 2009a). PMF decomposes a matrix of ambient data into two matrices, which an analyst then interprets to identify the represented source types. The method is described in greater detail elsewhere (Paatero, 1997; Paatero and Tapper, 1994). Ulbrich et al., and others (Lanz et al., 2008a, 2007) have also

expanded on the details of PMF application to AMS data, including a PMF analysis package in Igor Pro (Ulbrich et al., 2009a). An ambient data set can be viewed as a data matrix  $X$ , in which rows (denoted by  $i$ ) correspond to samples and columns (denoted by  $j$ ) corresponding to chemical species or, in the case of AMS data, to fragments of different  $m/z$ . The goal is to reduce the data set to a small number of factors that best characterize the OM composition with a profile  $f$  of each factor and a contribution  $g$  for each factor to each sample, plus residuals  $e$ .

EPA PMF, a freely available data analysis software package that utilizes the multilinear engine (ME-2) to solve the PMF equations described above, was used in this application (Norris et al., 2008, 2009). New features within EPA PMF v4.0 allow the user to take advantage of rotational tools available in ME-2 (Paatero, 2004). In general, the non-negativity constraint alone in PMF analysis is not always sufficient to produce a unique solution. To reduce the number of solutions, additional information such as known source contributions and/or source compositions can be used. This additional information can be incorporated into the PMF solution by “pulling” parts of a PMF solution, such as a factor profile or contribution. For example, if a source type has a typical ratio among elements in its source profile, a PMF-resolved factor profile could be pulled toward that ratio if the user has good confidence that the factor is related to such a source. The strength of each pull is controlled by specifying a limit on the change in the goodness-of-fit parameter  $Q$ ,  $dQ$ . If the user wishes to perform a weak pull, a small limit on  $dQ$  would be allowed. For a stronger pull, a large limit  $dQ$  would be allowed. These pulls are activated in ME-2 by the use of an optional control file called “moreparams.txt”, which is generated by EPA PMF when a user specifies a pull, or can be generated as a text file by a user outside of EPA PMF and ME-2. When the moreparams file is present, ME-2 generates a solution where the base solution is pulled as indicated in the moreparams file. Additional details are available in Paatero (2004) and Norris et al. (2009).

### 2.4 AMS data for EPA PMF application

A matrix of 7455 2-min HR-AMS V-mode observations of 198 unit mass resolution (UMR) fragments made every 4 min during January 2008 at Fyfe was used in EPA PMF analysis. Fragments predominantly from inorganic species such as nitrate and sulfate were not retained for EPA PMF analysis. Fragments  $m/z$  15, 18, 19, 20, 29, and 30 were excluded from EPA PMF analysis because of potential interference with nitrogen and oxygen. While data up to  $m/z$  700 are available, many fragments above  $m/z$  200 had low signal-to-noise (S/N) ratios (i.e., less than 6), made a minimal contribution to total OM, and were collinear with a number of other fragments. Fragments up to  $m/z$  240 were retained for EPA PMF, a total of 198 fragments. Fragments with low S/N may bias the results (Norris et al., 2008), so the uncertainties of a

given fragment were multiplied by 3 if the S/N for the fragment was less than 6; this reduces the fragment's influence on the solution. A global 10 % uncertainty was also applied to account for additional modeling uncertainty (Norris et al., 2008). EPA PMF was run in the robust mode, which reduces the influence of outliers.

Each observation was also classified as downwind (wind speed greater than  $2 \text{ m s}^{-1}$  and wind direction between 90 and 270 degrees;  $N = 1360$ ); upwind (wind speed greater than  $2 \text{ m s}^{-1}$  and wind direction between 310 and 60 degrees;  $N = 949$ ); other (wind speed greater than  $2 \text{ m s}^{-1}$  and wind direction between 60 and 90 degrees or between 270 and 310 degrees;  $N = 461$ ); or calm (wind speed less than  $2 \text{ m s}^{-1}$ ;  $N = 4907$ ). As an additional set of runs, EPA PMF was also applied to downwind-only data to examine whether factor profiles change, and how factor contributions change. Since downwind-only data are more heavily influenced by the freeway, we may expect that a downwind-only HOA factor would be more similar to vehicle exhaust source profiles than when all data are used.

The AMS PMF results were averaged up to 20-min intervals to allow for matching with the collocated 5-min data, which were also averaged up to 20-min intervals. The 20-min averaged AMS data ( $N = 1491$ ) were also re-analyzed with EPA PMF to evaluate the impact of high-mass transient events; profiles and contributions were then compared to the results using 2-min data. Factor profiles were compared to PMF factor profiles from earlier ambient studies (Lanz et al., 2008a, 2007; Ulbrich et al., 2009a) as well as to source profiles (Mohr et al., 2009; Sage et al., 2008; Weimer et al., 2008) originating from a publicly available online database of reference spectra hosted by University of Colorado (Ulbrich et al., 2009a, 2009b). Specifically, source profiles of Honda gasoline exhaust and diesel exhaust (Mohr et al., 2009), PMF-resolved OOA and HOA factor profiles from Pittsburgh (Zhang et al., 2005a), BBOA factor profiles from Switzerland (Lanz et al., 2007, 2008b), aged diesel exhaust profile from chamber experiments (Sage et al., 2008), and burning and smoldering oak and chestnut wood source profiles (Weimer et al., 2008) were used.

### 3 Results

#### 3.1 Ambient aerosol variability and composition

OM averaged  $3.3 \mu\text{g m}^{-3}$  during the January intensive campaign and was typically highest during the evening hours (i.e., 19:00 through 21:00 LST) with a secondary peak in the morning (i.e., 06:00–09:00 LST, during rush hour commute) as shown in Fig. 2. AMS fragments associated with HOA, such as  $m/z$  57 and  $m/z$  43, showed a similar diurnal pattern. AMS fragments associated with OOA, such as  $m/z$  44 ( $\text{COO}^+$ ), showed only minor fluctuations throughout the day, while those used as tracers of biomass burning, such as  $m/z$

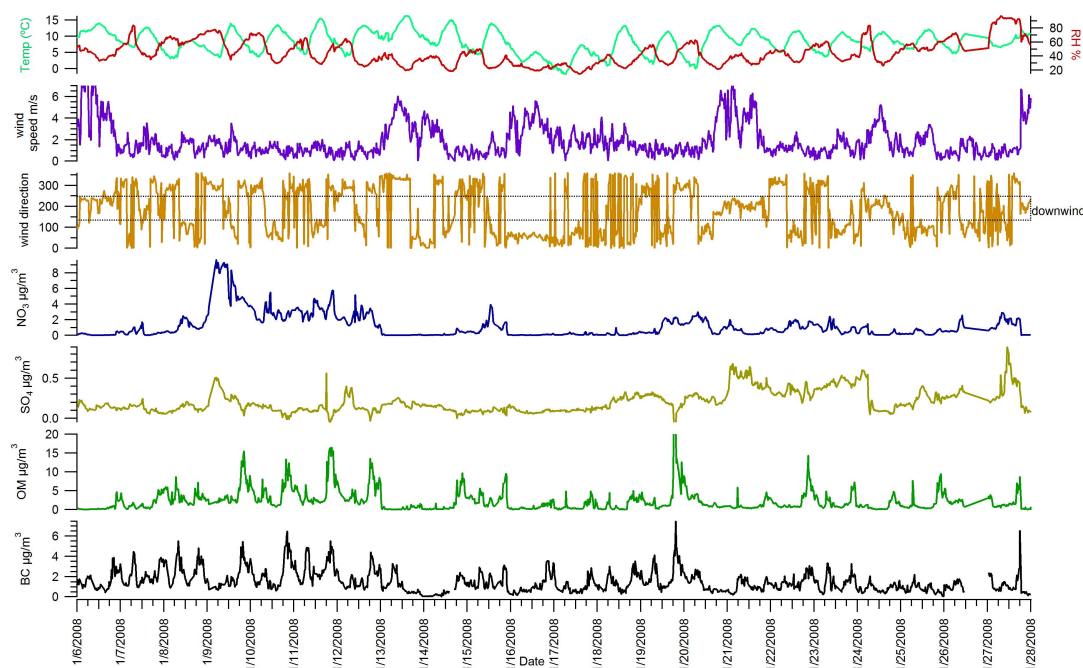
60 (Lanz et al., 2007, 2008b), were evident only during the evening and overnight hours.

Concentrations of other species such as BC, CO, and  $\text{NO}_x$  showed peaks in the early morning and overnight. There was only modest correlation (i.e.,  $r^2$  of 0.40 to 0.60) of OM and fragments such as  $m/z$  43, 44, 57, and 60 with BC, CO, and  $\text{NO}_x$ . Since BC and CO are from primary emissions and OM is a mixture of primary emissions and secondary products of atmospheric reactions, the modest correlations among these parameters suggest that a large fraction of the OM may be secondary in nature. OM concentrations were similar between upwind and downwind conditions (medians of  $2.0 \mu\text{g m}^{-3}$  and  $1.6 \mu\text{g m}^{-3}$ , respectively); there was a higher median OM concentration during upwind conditions since these occurred during the evening, when wind speeds were lower and the boundary layer shallower than during the daytime, when downwind conditions prevailed and higher wind speeds and more dispersion occurred. In contrast, other pollutants showed sharper differences between upwind and downwind conditions; for example, concentrations of primary emission pollutants such as BC were more than three times higher under downwind conditions than under upwind conditions (Roberts et al., 2010). Rather than being highest during downwind conditions, OM was highest during stagnant, low-wind conditions, with an average of  $3.7 \mu\text{g m}^{-3}$ .

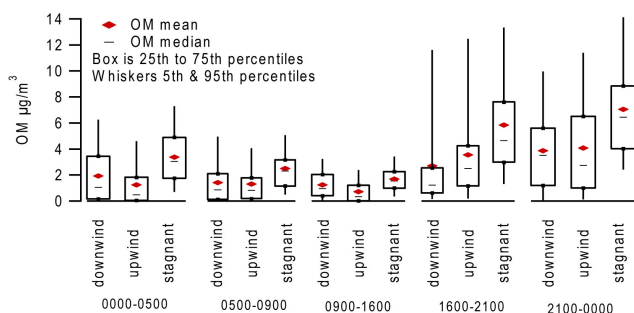
While OM did have a distinctive diurnal pattern in general, it was episodic during the intensive. A multi-day OM episode occurred with relatively high, sustained concentrations in the first week (Fig. 3). The episode ended around midnight on 12 January when a storm front came through the area. During the episode, OM concentrations were relatively high during the overnight periods under a stable boundary layer. A short-duration but very high OM concentration episode occurred the evening of 19 January, with the highest 2-min and hourly averaged OM concentrations of the entire intensive. Although sampling took place over four weekends, the meteorological conditions during those weekends were quite different, with drizzle and rain on the first weekend (Sunday, 6 January), a front with high winds and rain on the second (Saturday night 12 January), a stagnation episode on the third (Saturday 19 January), and windy conditions on the last weekend. With the low number of weekends and the wide range of meteorological conditions, comparing weekday to weekend concentrations may not be as useful here compared to data sets that comprise many weekends.

OM concentrations were generally similar whether our monitoring site was upwind or downwind of the freeway (Fig. 4). This is different than for other pollutants such as BC, where downwind concentrations were, on average, more than two times higher during downwind conditions. This shows that the enhancement of being next to the roadway is not nearly as large for OM as for BC, since OM is a mixture of primary, semi-volatile, and more oxidized material, whereas BC is primary material that, in a near-roadway environment, predominantly originates from emissions along





**Fig. 3.** Time series of temperature (degrees C), relative humidity (%), wind speed ( $\text{m s}^{-1}$ ), wind direction, AMS nitrate ( $\mu\text{g m}^{-3}$ ), AMS sulfate ( $\mu\text{g m}^{-3}$ ), AMS OM ( $\mu\text{g m}^{-3}$ ), and Aethalometer BC ( $\mu\text{g m}^{-3}$ ). Major tick marks indicate midnight for each day. Periods during which the monitoring site is downwind of the freeway are indicated by the dashed rectangle on the wind direction time series.



**Fig. 4.** Box plot of OM concentrations ( $\mu\text{g m}^{-3}$ ) during downwind, upwind and stagnant (wind speed less than  $0.5 \text{ m s}^{-1}$ ) conditions, grouped by time of day.

the roadway. Like BC, OM was higher during stagnant conditions (those with winds less than  $0.5 \text{ m s}^{-1}$ ), as a shallow boundary layer and minimal dispersion quickly led to the buildup of pollutants. During the daytime when emissions from the freeway may be most prominent, concentrations during downwind conditions were higher than during upwind conditions. Only during the evening hours (17:00–23:00 LST), when non-mobile emission sources become important, were upwind concentrations higher than downwind. This increase is likely due to contributions from residential biomass burning, as explored further with PMF.

Temporal patterns for both sulfate and nitrate differed from the temporal pattern for OM. Sulfate concentrations were extremely low throughout the study, with a median concentration of  $0.16 \mu\text{g m}^{-3}$ , and a maximum 20-min average of  $0.88 \mu\text{g m}^{-3}$ . Nitrate concentrations were episodic but were also typically low, with a median concentration of  $0.54 \mu\text{g m}^{-3}$ . Nitrate had little relationship with OM or other measured pollutants. In a one-day episode, nitrate peaked above  $8 \mu\text{g m}^{-3}$  for three hours, after which it decreased to approximately  $2 \mu\text{g m}^{-3}$  for the next two days. This episode occurred after a 24-h period when winds were stagnant, temperatures were low (less than  $8^\circ\text{C}$ ), and relative humidity was relatively high (greater than 55 %), including during the midday, which was atypical. Nitrate was greater than  $1 \mu\text{g m}^{-3}$  on a few other days, but these periods typically lasted only a few hours. These higher concentrations occurred during the day and night, with no distinct, consistent diurnal pattern, unlike OM, BC, CO, and other pollutants, which peaked in the evening and early morning hours. In contrast to the inorganic species, BC had a median concentration of  $1.24 \mu\text{g m}^{-3}$ , more than twice as high as nitrate, an order of magnitude higher than sulfate, and a third of OM. BC is more than twice as high at this site compared to a site 2 km away in the urban center but away from freeways (Hancock Elementary School), where BC was, on average,  $0.5 \mu\text{g m}^{-3}$ .

These sulfate levels are quite low compared to most other areas in the US, though typical of the western US, as there

is minimal transport of sulfate into Las Vegas and no major sources of  $\text{SO}_2$  or sulfate upwind of the urban area. The nitrate concentrations are also lower than is typical in western urban areas. There have been limited studies of speciated  $\text{PM}_{2.5}$  in Las Vegas, with the 2000–2001 Las Vegas Valley Visibility/ $\text{PM}_{2.5}$  study the most detailed and recent. Here they found wintertime sulfate concentrations, on average, lower than  $0.5 \mu\text{g m}^{-3}$ , and average nitrate concentrations between  $0.2$  and  $0.6 \mu\text{g m}^{-3}$  at three sites (Green et al., 2002). There are very few sources upwind, so the amount of transported secondary organic carbon and ammonium sulfate is generally low, though there can be transport from California. Some speciated  $\text{PM}_{2.5}$  data have been collected as part of EPA's chemical speciation network (CSN) from 2002 to 2007. Using the data from this 5-yr period, average sulfate concentrations are  $1 \mu\text{g m}^{-3}$ , and average nitrate concentrations are  $0.89 \mu\text{g m}^{-3}$ , though sulfate is lower and nitrate is higher in the winter. These low values are corroborated by Chapter 2 of the IMPROVE network annual report (Hand et al., 2011), which shows that sulfate and nitrate concentrations are extremely low in Las Vegas compared to other areas in the US and are lower than  $1 \mu\text{g m}^{-3}$ , on average. Unlike most of the US or much of Europe, there is very little coal-fired power production upwind of Las Vegas, so  $\text{SO}_2$  (and sulfate) levels are low. In many prior studies elsewhere, sulfate concentrations are higher and are correlated with LV-OOA concentrations; however, as the sulfate levels are so low in Las Vegas, it is unlikely that LV-OOA will correlate with sulfate. As Las Vegas is in an arid desert environment with little agriculture in the area, ammonia emissions are relatively low (e.g., <http://nadp.sws.uiuc.edu/nh3net/>), so ammonium nitrate formation may be limited by the availability of ammonia. Nitrate has been observed to often correlate with SV-OOA, as both may be associated with similar transport and formation mechanisms. When nitrate formation is limited by ammonia availability, however, there is little reason to expect it to correlate with SV-OOA.

### 3.2 EPA PMF application to AMS data summary

Three- to six-factor solutions were explored with EPA PMF. Initially, 50 runs from a random seed were performed for each number of factors. Random starting seeds were used to increase the likelihood of finding a global minimum of the goodness-of-fit parameter,  $Q$ . The stability of  $Q$  over these runs, the ratio of  $Q$  to expected (theoretical)  $Q$ , scaled residuals, the  $Q/Q_{\text{expected}}$  by fragment and sample, and factor independence (G-space plots) were examined. If these parameters are not stable for a given number of factors, it indicates that a global minimum was not consistently achieved, and that a solution may not be stable (Norris et al., 2009, 2008). In all solutions, the total OM was well apportioned (i.e., slope equal to  $1.0 \pm 0.10$ ) and  $r^2$  was greater than 0.95 between apportioned and total OM.

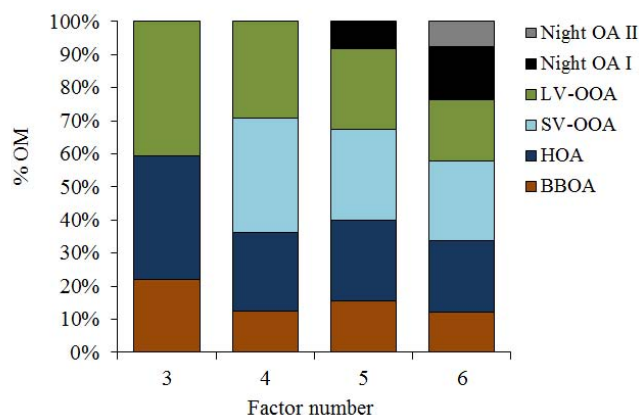


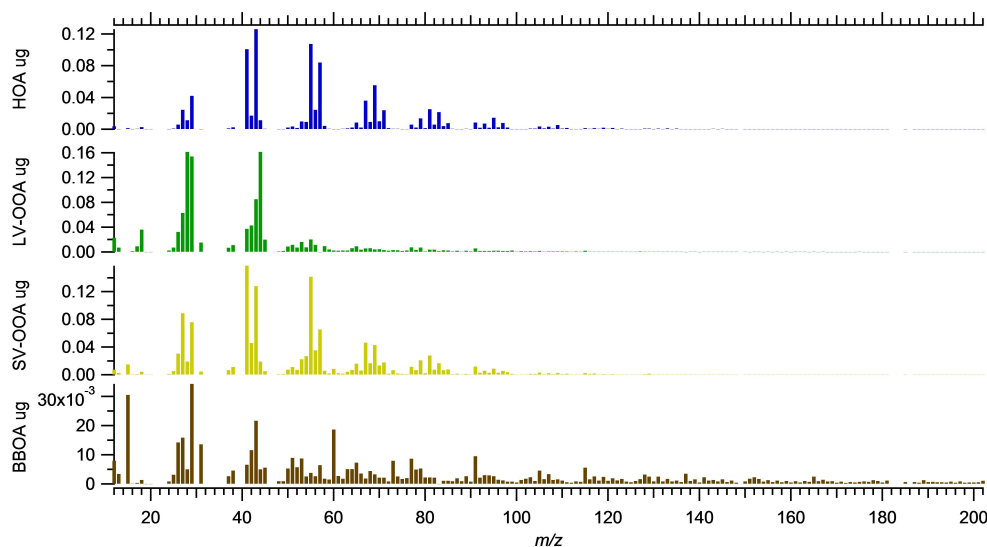
Fig. 5. Summary of OM apportioned by factor number.

PMF factors were classified by their temporal pattern and the comparison of their profile with source profiles and profiles from previous studies. The typical AMS PMF factors of HOA, LV-OOA, and BBOA were resolved in every solution with three or more factors, and the identification of each factor profile was based on its similarity to profiles available in the literature, the abundance of key fragments in each profile, and each factor's temporal pattern. For example, the LV-OOA factors displayed a significant amount of  $m/z$  44 and were similar to the OOA factor identified in Pittsburgh and elsewhere. BBOA factors had typical tracer fragments of  $m/z$  60 and 73, which are produced during AMS analysis of levoglucosan and related anhydrosugars produced during biomass combustion (Lanz et al., 2008b, 2010; Alfarrá et al., 2004). The HOA factors were similar to the HOA factor from Pittsburgh and a diesel exhaust source profile.

With more than three factors, semi-volatile OOA (SV-OOA) was also resolved. This factor's profile was similar to that of aged diesel exhaust and was more episodic than HOA or LV-OOA profiles. With five and six factors, additional factors that occurred during the nighttime were resolved. The additional nighttime factors occurred nearly every night coincident with BBOA between 17:00 and 02:00 LST, and contributions were higher with low wind speed conditions and with winds from the north (upwind). With peaks of  $m/z$  41, 43, 55, and 91, it is unclear what these "night OA" factors may represent. The change in the BBOA between the four- and five-factor solutions and its temporal variability suggest the additional fifth factor may be related to biomass burning, but the lack of known tracer fragments and correlation with source profiles makes this link difficult to prove. We focused the remainder of the analyses on the four-factor solution. Figure 5 summarizes the amount of OM apportioned by factor. Table 1 summarizes the correlation of factor profiles with selected source, aged source, and PMF factor profiles from other studies.

**Table 1.** Correlation ( $r^2$ ) of PMF factor profiles with Pittsburgh<sup>1</sup> OOA PMF profile, Pittsburgh HOA PMF profile, diesel exhaust source profile<sup>2</sup>, aged diesel exhaust profile<sup>3</sup>, charbroil<sup>4</sup> source profile, oak-flame<sup>5</sup> source profile, oak-smolder source profile, chestnut-flame source profile, chestnut-smolder source profile, and levoglucosan<sup>6</sup> profile. Correlations from 0.80 to 0.90 are denoted in *italics*, and those greater than 0.90 are denoted in **bold**.

N	Factor	Pitt HOA	gas- oline	Diesel OOA	Pitt diesel	Aged	charbroil smolder	Oak flame	Oak glucosan	levo- flame	Chestnut smolder	Chestnut
3	LV-OOA	0.59	0.58	0.58	<b>0.98</b>	<b>0.95</b>	0.26	0.88	<b>0.91</b>	0.74	<i>0.83</i>	<i>0.89</i>
	HOA	<b>0.97</b>	<b>0.92</b>	<b>0.95</b>	0.46	0.75	0.57	0.29	0.71	0.49	0.50	0.62
	BBOA	<i>0.80</i>	0.73	0.76	0.67	<i>0.82</i>	0.40	<i>0.44</i>	<i>0.89</i>	<i>0.77</i>	<i>0.77</i>	<i>0.85</i>
4	LV-OOA	0.38	0.39	0.38	<b>0.99</b>	0.87	0.13	<b>0.93</b>	0.83	0.70	0.80	0.84
	HOA	<b>0.99</b>	<b>0.96</b>	<b>0.98</b>	0.48	0.77	0.60	0.32	0.69	0.48	0.50	0.62
	BBOA	0.47	0.41	0.44	0.59	0.64	0.20	0.39	0.75	<i>0.80</i>	<i>0.77</i>	<i>0.80</i>
5	SV-OOA	<b>0.91</b>	<i>0.84</i>	<i>0.87</i>	0.57	0.79	0.49	0.37	0.80	0.58	0.60	0.71
	LV-OOA	0.46	0.46	0.46	<b>0.99</b>	<b>0.90</b>	0.18	<b>0.91</b>	0.88	0.72	<i>0.82</i>	<i>0.87</i>
	HOA	<b>0.99</b>	<b>0.97</b>	<b>0.98</b>	0.47	0.78	0.61	0.32	0.68	0.49	0.50	0.62
6	BBOA	0.71	0.69	0.69	<i>0.86</i>	<b>0.91</b>	0.37	0.69	<b>0.95</b>	<i>0.89</i>	<b>0.91</b>	<b>0.97</b>
	SV-OOA	<i>0.90</i>	<i>0.82</i>	<i>0.85</i>	0.56	0.78	0.47	0.37	0.79	0.56	0.58	0.69
	Night OA I	0.60	0.51	0.59	0.42	0.61	0.27	0.26	0.57	0.30	0.35	0.45
6	LV-OOA	0.13	0.16	0.15	0.59	0.46	0.05	0.54	0.47	0.44	0.51	0.51
	HOA	0.12	0.17	0.14	0.55	0.38	0.11	0.60	0.42	0.43	0.49	0.48
	BBOA	0.24	0.26	0.24	0.72	0.57	0.13	0.66	0.59	0.56	0.63	0.64
6	SV-OOA	0.24	0.28	0.25	0.61	0.48	0.12	0.60	0.54	0.54	0.59	0.61
	Night OA I	0.33	0.36	0.34	0.54	0.48	0.16	0.47	0.53	0.55	0.56	0.61
	Night OA II	0.14	0.16	0.15	0.49	0.35	0.03	0.47	0.42	0.42	0.44	0.47



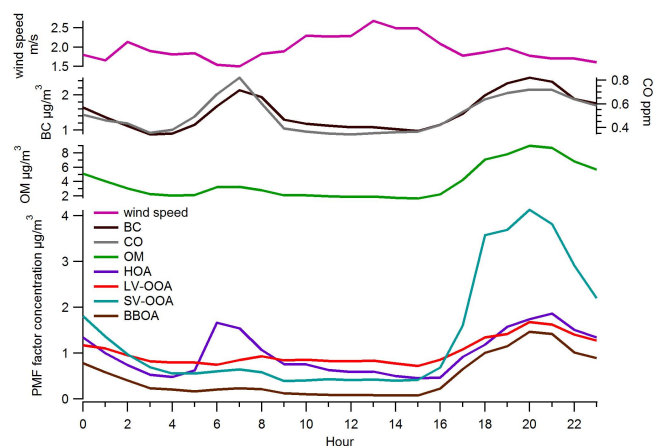
**Fig. 6.** PMF factor profiles through  $m/z$  200 for the four-factor solution.

### 3.3 Four-factor solution

The factors in the four-factor solution were HOA, LV-OOA, and BBOA factors, plus a semi-volatile OOA (SV-OOA) factor. The HOA and LV-OOA factors were better resolved than in the three-factor solution. Profiles of each factor, the average factor concentration plus other species' concentra-

tions by hour, and a time series of concentrations are provided in Figs. 6 through 8. Figure 9 provides scatter plots of factor contributions with selected collocated measurements: HOA with BC, HOA with CO, SV-OOA with nitrate, and LV-OOA with sulfate. Factor profiles were compared to profiles available in the literature to help confirm identification; regression statistics were reported using Pearson correlation





**Fig. 7.** PMF factor contributions, OM, BC, CO, and wind speed averaged by hour.

values, and each scatter plot associated with the regression statistics was examined to ensure that the correlation was not biased by the large range of concentrations of the individual  $m/z$  fragments. Bootstrapping, in which many runs are used to gauge the uncertainty of the base solution (in this case 300 runs with an  $r^2$  of 0.60), showed good reproducibility of the factors. All factors were reproduced at least 98 % of the time, demonstrating that these factors are stable and characterize the solution space well.

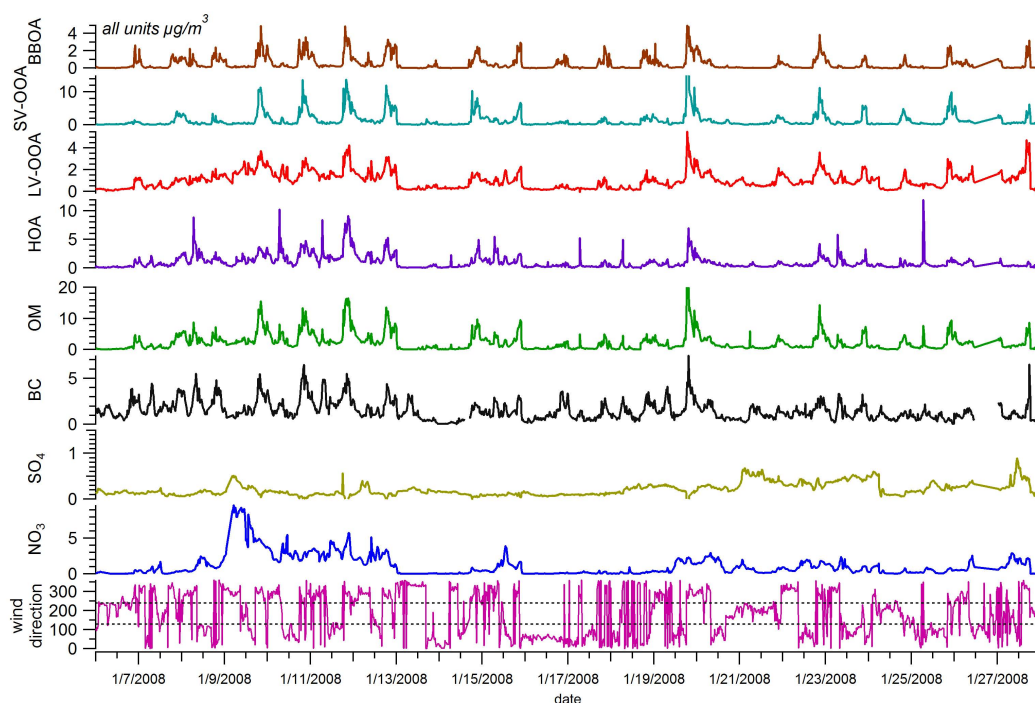
The LV-OOA factor displayed the typical high amount of  $m/z$  44 but with a lower amount of  $m/z$  43 than in the three-factor solution, and it showed a high correlation with the Pittsburgh OOA factor profile ( $r^2$  of 0.99; Table 1). LV-OOA accounted for 26 % of the OM with four factors and showed only small diurnal variability in its concentration, though it was a much greater percentage of the OM during the day-time than at other times. Similar to other factors and total OM, LV-OOA concentrations were lower at higher wind speeds, though it was a higher percentage of OM at higher wind speeds. LV-OOA showed little correlation with other pollutants; ozone, which often showed a moderate correlation with LV-OOA in other studies, was not measured here. Sulfate is also often correlated with LV-OOA, but during this study sulfate levels were extremely low, with a median of  $0.16 \mu\text{g m}^{-3}$ . LV-OOA is heavily oxidized and likely part of a background OM, and it may be transported into Las Vegas over multiple days. In addition, there are very few  $\text{SO}_2$  sources upwind of Las Vegas, so there is very little sulfate transported into the area. Thus, we may not expect LV-OOA concentrations to be correlated with sulfate here.

The BBOA factor accounted for 12 % of OM, on average, and was similar to the BBOA factor found in the three-factor solution. This factor had more than 60 % of the  $m/z$  60 fragment, which is associated with levoglucosan and related anhydrosugars (Lanz et al., 2008b; Alfara et al., 2004); the contribution from BBOA was well correlated with  $m/z$  60

( $r^2 = 0.86$ ). This factor is most likely from residential wood burning in the evenings rather than wildfire emissions since there was little regional wildfire activity in the winter; furthermore, the contribution was nearly zero during the day-time, with a sharp rise in concentrations nearly every evening after 17:00 LST that peaked around 21:00 LST. While the BBOA factor concentration began to decrease after around 21:00 LST, its relative contribution to OM remained above 15 % until after midnight. The factor was highest in terms of both concentration and relative contribution to OM under low wind speed conditions (i.e., less than  $2 \text{ m s}^{-1}$ ). This profile is similar to a smoldering Chestnut profile ( $r^2 = 0.80$ ) and the levoglucosan combustion profile ( $r^2 = 0.80$ ) (Schneider et al., 2006). This BBOA factor is mostly associated with winds from the north and west—the direction of a large residential neighborhood upwind of the freeway. When using five factors, the BBOA factor profile has even better correlations with both of these profiles ( $r^2$  of 0.89 and 0.91), plus an  $r^2$  of 0.95 with an oak flaming profile. However, since the fifth factor is unidentified, we have focused on the four-factor solution.

The HOA factor accounted for 26 % of the OM and had peaks of  $m/z$  41, 43, 55, 57, and other fragments typical of hydrocarbon-like fragments. HOA concentrations were highest during the early morning and overnight periods, and as a percent of total OM, the factor's contributions were highest during the early morning (06:00–08:00 LST). The HOA factor profile has a high correlation with diesel exhaust ( $r^2 = 0.98$ ) and gasoline exhaust ( $r^2 = 0.96$ ) source profiles (Mohr et al., 2009). The HOA profile here was very similar to that observed in Pittsburgh ( $r^2 = 0.99$ ). This factor is likely heavily influenced by the mobile emissions on the adjacent freeway but may also have originated in part from other sources. Similar to concentrations of BBOA and other pollutants such as BC, HOA concentrations rapidly decreased with increases in wind speed, though on a relative basis there was no significant difference in its contribution. With sustained winds (i.e., greater than  $2 \text{ m s}^{-1}$ ), HOA concentrations were significantly higher under downwind conditions. HOA had modest correlation with collocated measurements of CO ( $r^2 = 0.66$ ),  $\text{NO}_x$  ( $r^2 = 0.64$ ), and BC ( $r^2 = 0.68$ ). This may be in part because BC, CO, and  $\text{NO}_x$  have large differences between upwind and downwind conditions. For example, during November–March, BC is twice as high under downwind conditions (average  $2.3 \mu\text{g m}^{-3}$ ) than upwind conditions (average  $1.2 \mu\text{g m}^{-3}$ ), while OM (and HOA) do not have as large a difference.

The last factor resolved was semi-volatile OOA (SV-OOA), which accounted for 35 % of the OM. It showed strong peaks of  $m/z$  41, 43, 55, 57, 67, 69, and 71 but also had some contribution from  $m/z$  44, with a 43/44 ratio of 6, almost half the value for HOA (11). The SV-OOA factor profile had a moderate correlation with Pittsburgh HOA and with chamber-aged diesel exhaust ( $r^2$  of 0.91 and 0.79, respectively). SV-OOA contributions were highest in the evening



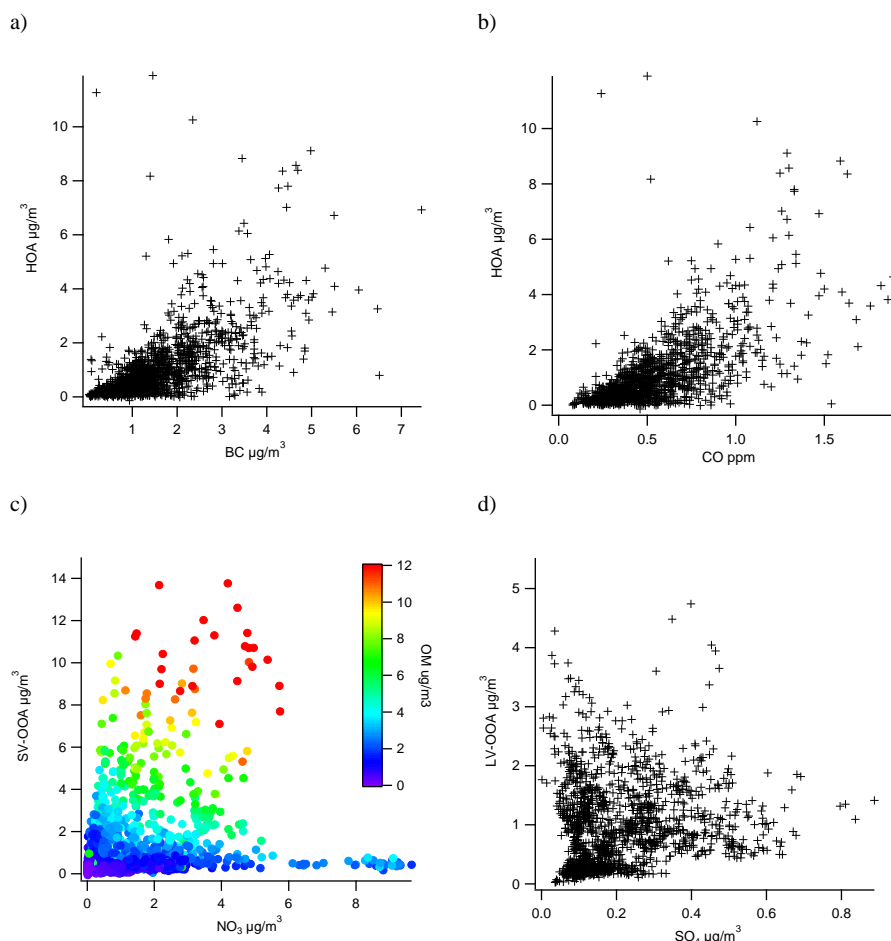
**Fig. 8.** Time series of PMF factor contributions, BC, sulfate, nitrate, and OM (all in  $\mu\text{g m}^{-3}$ ), plus wind direction. Times when the monitoring site is downwind are outlined by the dashed box.

and overnight hours, though on a relative basis its contributions were generally very consistent across all hours. Like HOA, SV-OOA factor contributions decreased with higher wind speeds, but its relative contribution was not significantly different among wind speed ranges. In other studies, this factor sometimes has a modest correlation with nitrate. Here, nitrate was extremely episodic, in that its concentrations were less than  $0.5 \mu\text{g m}^{-3}$  half the time, with a day-long episode of concentrations greater than  $4 \mu\text{g m}^{-3}$  and a few hours during which concentrations intermittently exceeded  $2 \mu\text{g m}^{-3}$ . In contrast, SV-OOA was present during nearly the entire study and had a modest diurnal pattern similar to other factors where it was highest in evening hours.

It is also possible that the SV-OOA factor contains some contribution from cooking organic aerosol (COA), as the SV-OOA factor shows similarities in both profile and temporal patterns to a COA factor found in London and Manchester (Allen et al., 2010). Similar to the COA profile, the largest peaks in the SV-OOA profile are at  $m/z$  41 and  $m/z$  55, and in both profiles  $m/z$  41 is greater than  $m/z$  43 and  $m/z$  55 is more than twice  $m/z$  57; this is in contrast to HOA, where  $m/z$  43 is greater than  $m/z$  41, and  $m/z$  55 is only slightly higher than  $m/z$  57. Similar to the COA factor, our SV-OOA factor peaks, on average, in the evening, again in contrast to HOA, which peaks in the morning and evening hours associated with rush hour. There are, however, times when the SV-OOA factor is high that are not likely periods of high cooking emissions, such as late morning or midday. Expanding to five

factors was inconclusive, since the fifth factor was similar to SV-OOA and COA but was generally collinear with BBOA; it was unclear what this fifth factor represented, so we have retained the SV-OOA label for this factor, with the caveat that there is likely some COA influence.

To help confirm factor identification and understand the OM composition difference between upwind and downwind conditions, we examined the factor contributions during downwind ( $N = 1360$ ) and upwind ( $N = 949$ ) conditions and compared them to the average over the study period. We also examined the OM composition under downwind conditions ( $N = 195$ ) from 05:00 to 09:00 LST, when the impact from emissions on the freeway is expected to be highest. Results are summarized in Fig. 10. As expected, HOA contributions are higher under downwind conditions; during morning downwind conditions, HOA accounts for 49 % of the OM. SV-OOA was, on average, similar during upwind, downwind, and stagnant conditions, except during morning downwind situations when it was only 23 % of the OM. The minimal difference with different wind directions suggests that SV-OOA is a slightly aged factor that is not characteristic of direct, primary emissions. BBOA contributions were low, on average, under downwind conditions and higher (16 % on average) during upwind conditions. LV-OOA contributions were relatively lower during morning downwind conditions and relatively highest during midday periods, regardless of wind direction.



**Fig. 9.** Scatter plot comparisons of PMF factor contributions for: (a) HOA and BC; (b) HOA and CO; (c) SV-OOA and nitrate; and (d) LV-OOA and sulfate.

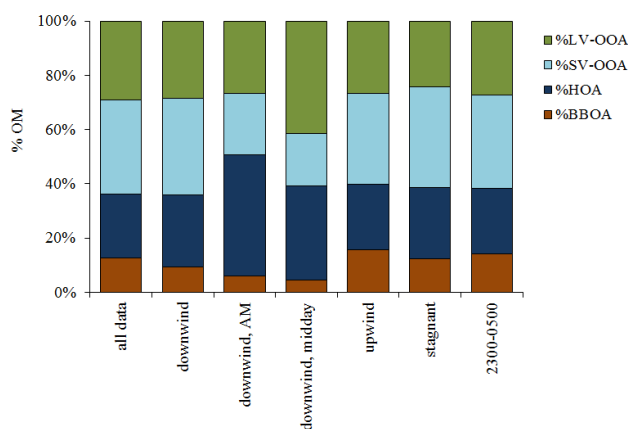
### 3.4 Further analysis using ME-2 rotational tools in EPA PMF

$F_{\text{peak}}$  is a parameter available in PMF to rotate the entire solution, a process that can help indicate if there is rotational freedom in the solution. G-space plots of the base solution show distinct edges, indicating some factor interdependence. To ascertain whether the solution changes or whether these edges can be rotated to the y- and x-axes, rotation using  $F_{\text{peak}}$  was performed. PMF runs were conducted with  $F_{\text{peak}}$  values at increments of 2 between 8 and  $-8$  for a total of eight runs. In general, minimal change was seen in the factor profile, contributions, and G-space plots;  $Q$  increased by less than 0.1 % for the runs with the highest  $F_{\text{peak}}$  value and by even less under other  $F_{\text{peak}}$  values. Since there is little change in the solution with  $F_{\text{peak}}$ -induced rotations, the base solution appears to be rotationally unique. The oblique edges in the G-space plots may be due to co-dependence among factors, or modeling errors, such as variation in true source profiles during the monitoring campaign.

The four-factor solution was further explored with ME-2 rotational tools available in EPA PMF. In each scenario, fragment ratios in factor profiles were pulled toward source profile ratios. In one scenario, the ratio of  $m/z$  43/44 in the SV-OOA factor (6.75) was pulled toward the  $m/z$  43/44 ratio in the 5-hour aged diesel exhaust profile (1.34). In another scenario, the  $m/z$  57/55 and 41/43 ratios in HOA (0.78 and 0.80) were pulled to the ratios in the diesel exhaust profile (1.03 and 0.69, respectively). The BBOA factor ratio of  $m/z$  60/91 was pulled to the chestnut smolder profile ratio (1.96 to 7.20). Lastly, 157 points along an apparent edge in the SV-OOA versus LV-OOA G-space plot were pulled in an attempt to force additional independence between these factors. Seeing how the solutions change, in terms of factor profiles,  $dQ$ , and G-space plots, can help us understand the stability of the original solutions. Different maximum  $dQ$  values were allowed for each combination; 1 % and 3 % of  $Q(\text{robust})$  were calculated, and these values were used for the maximum  $dQ$  allowed for a given pull. In the G-space pull, each point was allowed a  $dQ$  of 0.2 %, for a total  $dQ$  of 31 %.

**Table 2.** Summary of pulls on the four-factor solutions, all with a maximum allowed  $dQ$  of 1 %, except for one iteration with BBOA to Chestnut smolder ( $dQ = 3$  %), and except for the edge points pull, which had a total allowed  $dQ$  of 31 %, or 0.2 % per point with a total of 157 points.

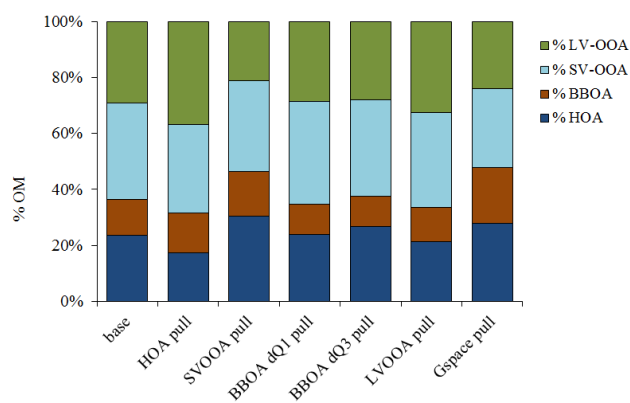
Pull	SV-OOA to Aged Diesel	HOA to Diesel	BBOA to Chestnut Smolder ( $dQ = 1$ %)	BBOA to Chestnut Smolder ( $dQ = 3$ %)	Edge points on SV-OOA vs. LV-OOA
Target	Ratio $m/z$ 43/44 from 6.75 to 1.34	Ratio $m/z$ 57/55 from 0.78 to 1.03 and ratio $m/z$ 41/13 from 0.80 to 0.69	Ratio $m/z$ 60/91 from 1.96 to 7.20	Ratio $m/z$ 60/91 from 1.96 to 7.20	157 points pulled to axis
Change in target values?	Yes	Yes	Yes	Yes	Yes
$dQ$	0.1 %	0.5 %	0.4 %	1.3 %	8.8 %
Improved G-space plot?	Worse LV-OOA vs. SV-OOA plot	No change	No change	No change	No change but for forced points
LV-OOA vs. Pittsburgh OOA	Declined $r^2$ from 0.99 to 0.93	No change	No change	No change	No change
HOA vs. Diesel	No change	Slightly better	No change	No change	No change
BBOA vs. Chestnut smolder	Slightly better; excl. $m/z$ 44	Improved $r^2$ from 0.79 to 0.87	Improved $r^2$ from 0.79 to 0.88	Improved $r^2$ from 0.79 to 0.86	Improved $r^2$ from 0.79 to 0.93
SV-OOA vs. Aged Diesel	Improved $r^2$ from 0.62 to 0.69	Slightly worse $r^2$ from 0.62 to 0.57	Worse $r^2$ 0.62 to 0.56, due to $m/z$ 44 = 0	Worse $r^2$ 0.62 to 0.56, due to $m/z$ 44 = 0	Improved $r^2$ from 0.69 to 0.72
Other changes of note	0 $m/z$ 44 in BBOA factor; unreasonable result	Large decrease in $m/z$ 44 in SV-OOA, resulting in 43/44 ratio of 21, higher than HOA ratio of 10	In SV-OOA $m/z$ 44 = 0; unreasonable result	In SV-OOA $m/z$ 44 = 0; in BBOA $m/z$ 55 = 0; unreasonable result	SV-OOA $m/z$ 43/44 ratio changed from 4.4 to 2.9



**Fig. 10.** Attribution of OM by factor in the four-factor solution over all data, during downwind conditions only, during downwind conditions between 05:00 and 09:00 LST only, during downwind conditions midday, during upwind conditions only, during stagnant conditions, and during 23:00–05:00 LST only.

The results from these pulls are summarized in Table 2 and Fig. 11. Since results were similar with  $dQ$  values of 1 % and 3 %, only the  $dQ$  of 1 % are shown, except for the BBOA pull, where both are provided as an example. In general, most pulls resulted in only minor changes in the HOA and LV-OOA factor profiles and contributions. For example, when the HOA factor was pulled toward the diesel exhaust profile, the correlation between the two increased by only 0.02. In all pulls, the correlation between BBOA and the chestnut smolder profile improved (e.g., correlation improved from 0.79 to 0.88 with the BBOA pull). However this was often at the expense of the SV-OOA profile, where the amount of  $m/z$  44 was much lower compared to the base solution, and in some pulls, was actually zero. While LV-OOA and HOA did not vary much between these pulled solutions, the changes in SV-OOA and BBOA profiles and contributions suggest some rotational freedom in these two factors. LV-OOA and HOA factors are similar across many studies, but the SV-OOA and BBOA factors vary among studies, and within this study under different pulling scenarios, because these factor profiles represent semi-fresh factors that are likely changing minute-to-minute in the atmosphere. Overall, the pulling results





**Fig. 11.** Attribution of OM with four factors for base run (base); ratio of  $m/z$  41/43 in the HOA profile pulled toward the diesel exhaust profile (HOA pull); ratio of  $m/z$  43/44 in the SV-OOA profile pulled toward the 5-h aged diesel profile (SV-OOA pull); ratio of  $m/z$  60/91 in the BBOA profile pulled toward the oak-flame source profile with different  $dQ$  values (BBOA  $dQ$  1 and 3 pull);  $m/z$  44 in LV-OOA profile pulled up maximally (LV-OOA pull); and edge points on the HOA/LV-OOA G-space plot pulled down to the y- and x-axes (Gspace pull).

indicate that the base solution is likely at a global  $Q_{\text{minima}}$ , and that there is little rotational freedom in the unpulled, base solution, in particular regarding the contributions of SV-OOA and BBOA.

With the G-space pull, 157 points along an edge in the LV-OOA/SV-OOA scatter plot were pulled to reduce LV-OOA to zero. LV-OOA concentrations on these points were successfully reduced towards zero, so that the edge was less well defined. This resulted in a  $dQ$  of 9 %, but an improvement in the comparison of the BBOA profile to the chestnut smolder profile ( $r^2$  from 0.79 to 0.93) and in the SV-OOA to aged diesel profile comparison ( $r^2$  from 0.69 to 0.72). The SV-OOA  $m/z$  43/44 ratio also decreased from 4.4 to 2.9, which is similar to the ratio of other studies summarized in Ng et al. (2010). HOA and BBOA contributions increased, while LV-OOA and SV-OOA contributions decreased. Even though  $Q$  increased by 9 %, these results are useful to show that contribution pulls based on the G-space plots helped improve the factor profiles, and to show what may be the bounds in the base solution results.

## 4 Discussion

Evaluating solutions with different numbers of factors, comparing profiles to source profiles, examining temporal trends, and exploring rotational ambiguity with the rotational tools available in ME-2 can lead to a greater understanding of the AMS data set. In general, all the factors were consistent under multiple scenarios, suggesting high confidence in their apportionment. Increasing the number of factors from four

helped to better characterize the solution space, but the additional “night OA” factors are not easily attributable to known sources. Oblique edges were present in G-space plots, suggesting that the obtained rotation may not be the most plausible one. Since solutions found by pulling the profiles or using  $F_{\text{peak}}$  retained these oblique edges, there appears to be little rotational freedom in the base solution.

The  $Q/Q_{\text{exp}}$  ratios for most fragments were around unity (i.e., between 0.8 and 1.3), indicating that the obtained  $Q$  values were approximately equal to the expected values (see Fig. 2 in Supplement for a summary). Some fragments had  $Q/Q_{\text{exp}}$  ratios below 0.5, indicating that computed  $Q$  values were significantly smaller than the expected  $Q$  values. This discrepancy is most likely due to the global uncertainty ( $s_{ij}$  increase of 10 %); since the  $Q/Q_{\text{exp}}$  ratios for some  $m/z$  were fairly low, the 10 % value may be too high for these  $m/z$  but appropriate for many other fragments. The  $Q/Q_{\text{exp}}$  ratios were between 1.3 and 2 for several  $m/z$ , indicating that the average residuals are between 11 and 14 % of  $x_{ij}$  for these  $m/z$ , since  $Q$  depends quadratically on the average size of the residuals. The  $Q$  contributions drop sharply beginning at  $m/z = 198$ , as most of these fragments had low S/N ratios and were downweighted. The large residuals may indicate that the PMF solution does not fully characterize these fragments<sup>7</sup>. Inaccurate subtraction of the inorganic component from  $x_{ij}$  before PMF analysis is a possibility, though many of the fragments with high  $Q/Q_{\text{exp}}$  do not have an inorganic component. More likely, the large residuals are due to the variation of factor profiles with time and/or to the presence of an occasional, spurious, or localized source(s).

While factors in the four-factor solution are similar to those observed elsewhere, the apportionment of mass among them is different from that seen in other studies. This difference is expected because a number of the previous studies occurred in the summer and/or in environments with a higher amount of oxidized aerosol than the present study. Previous studies in Los Angeles, Pittsburgh, rural British Columbia, and elsewhere have typically found that at least one-third of the OM was attributable to LV-OOA, originally termed OOA I (Allan et al., 2003a; Alfarra et al., 2004; Dzepina et al., 2007; Sun et al., 2009). The lower amount of LV-OOA observed during the wintertime Las Vegas study could be due to less transported/aged aerosol, lower biogenic emissions, and/or less overall atmospheric oxidation compared to summertime. In Zurich, a wintertime study found 52 % to 57 % of the OM to be LV-OOA (Lanz et al., 2008a), 69 % of which originated from non-fossil sources such as wood burning. In Las Vegas, there is a much lower concentration of BBOA than in Zurich, leading to a smaller concentration of LV-OOA from non-fossil sources and a smaller concentration of LV-OOA overall.

<sup>7</sup> Fragments with high  $Q/Q_{\text{exp}}$  include  $m/z$  = [44, 60, 73, 85, 86, 111, 112, 113, 114, 123, 124, 125, 126, 137, 138, 140, 141, 154, 155, 156].



## 5 Conclusions

EPA PMF v4.0, with its new rotational tools, was successfully applied to a near-road, high time resolution AMS data set. HOA was a quarter of the OM (24 %), and higher under downwind conditions (about 40 %). In addition to this local, primary OA, there was a highly oxidized background of OA (LV-OOA) that, on average, constituted 29 % of the OM, and a less oxidized, semi-volatile fraction that accounted for 34 % of the OM. During the evening hours, biomass burning (BBOA) was also seen, likely from the surrounding residential area. Rotational tools allowed for additional analysis of the PMF solution space, increasing our confidence in the results.

**Supplementary material related to this article is available online at:**

**<http://www.atmos-chem-phys.net/12/309/2012/acp-12-309-2012-supplement.pdf>**

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